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# Vacuum Energy of the Laplacian on the Spheres

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Author's contribution

The sole author designed, analyzed and interpreted and prepared the manuscript.

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## Abstract

Let  $\Delta_g$  be the Laplacian on smooth functions on a compact Riemannian manifold (M, g) and  $\zeta_g$  the associated spectral zeta function. Some special values of the spectral zeta function and their generalisations such as the spectral height and spectral determinant usually defined in terms of the spectral zeta function to be  $\zeta'_g(0)$  and  $\exp(\zeta'_g(0))$  respectively, have been computed explicitly, see e.g [1,2] and [3]. Another special value of the spectral zeta function which has been a fundamental issue in quantum field theory is the Vacuum (Casimir) energy. Casimir energy is defined, mathematically, via the spectral zeta function as a function on the set of metrics on the manifold by  $\zeta_g(-\frac{1}{2})$ , [4,5] and [6]. In

this paper, a general technique for computing the Casimir energy of the Laplacian on the unit n-dimensional sphere,  $S^n$ , by factoring the spectral zeta function through the Riemann zeta function  $\zeta_R$  is presented.

Keywords: Laplacian; spectral zeta function; Riemann zeta function; Casimir energy.

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## **1** Introduction

The study of Vacuum energy (also known as Casimir energy) is believed to originate from the work of Hendrik B. G. Casimir (1909 - 2000), who in the year 1948 pointed out the existence of a force between a pair of neutral perfectly conducting parallel plates, [7] and [8]. The Casimir energy may be thought-of as the energy difference due to the distortion of the vacuum, [6]. The energy difference gives rise to what is known as the Casimir force. Although Casimir energy is a concept arising in quantum field theory with observable consequences in physics, research is on-going in the modern aspects of spectral geometry, formulating the notion in a purely mathematical framework; you may see again [8] and the numerous literature cited there-in.

In most physics literature, the Casimir energy, popularly denoted by  $E_{cas}$  is written as sum over the eigenvalues  $\omega_k = \sqrt{\lambda_k}$  of the Laplacian on smooth functions on M, i.e.  $E_{cas} = \sum_k \omega_k$  which is the spectral zeta function  $\zeta_g$  at  $s = -\frac{1}{2}$ . Because the spectral zeta function depends on the choice of metric g, [9,10,4,11,12], the Casimir energy  $\zeta_g(-\frac{1}{2})$  is defined via the spectral zeta function as a function on the set of metrics on the manifold M, see [7]. However, this sum is usually divergent and has to be regularized.

The regularization of this a priori divergent sum has been variously addressed using different methods; see e.g [8,13,14,5,2,15,16,17,9] and mostly recently [3] and [6] among many other literature. For example in [7], Elizalde had to split the sphere into hemispheres and imposing Dirichlet boundary condition on one of the hemispheres and Neumann boundary condition on the other to be able compute Casimir energy on  $S^n$ , n = 1,2,3,4. The authors in [3] followed similar procedure as in [8] to compute functional determinant corresponding to massive Laplacian in arbitrary dimensional spheres. In this paper, a more general method of computing the Casimir energy  $\zeta_g(-\frac{1}{2})$  of the Laplacian  $\Delta_g$  on the *n*-sphere,  $S^n$  which only employs factoring the spectral zeta function  $\zeta_g(s)$  through the Riemann zeta function  $\zeta_R(s)$  is introduced. This method has an advantage of computing the Casimir energy of arbitrary dimensional spheres less tediously over the method of taking the full sphere as the union of Dirichlet and Neumann problems on hemispheres. This is simply obtained by computing the finite part, FP, of the spectral zeta function,  $\zeta_{S^n}$ , of the Laplacian on the spheres given by

$$\zeta_{s^n}(s) = 2\sum_{m=0}^{\infty} \sum_{k=1}^{\infty} \frac{(n-1)^{2m}}{m! 4^m} \frac{\Gamma(s+m)}{\Gamma(s)} \frac{\Gamma(k+n-1)}{\Gamma(n)\Gamma(k+1)} (k + \frac{n-1}{2})^{-2s-2m+1}.$$

for each n where the finite part function, FP, is defined by

$$FP[f](s) := \begin{cases} f(s) \text{ if } s \text{ is not a pole} \\ \lim_{\varepsilon \to 0} (f(s+\varepsilon) - \frac{\text{Residue}}{\varepsilon}), \text{ if } s \text{ is a pole.} \end{cases}$$

We proceed by giving the notion of the Laplacian and the spectral zeta function of the Laplacian on Riemannian manifolds.

## **2 Basic Concepts**

#### 2.1 The Laplacian on the unit *n*-sphere

The Laplacian on smooth functions on (M, g) is the operator

$$\Delta_{g}: C^{\infty}(M) \to C^{\infty}(M) \tag{2.1}$$

defined in local coordinates by

$$\Delta_{g} = -\operatorname{div}(\operatorname{grad}) = -\frac{1}{\sqrt{|g|}} \sum_{i,j} \frac{\partial}{\partial x^{i}} \left( \sqrt{|g|} g^{ij} \frac{\partial}{\partial x^{j}} \right)$$
(2.2)

where  $g^{ij}$  are the components of the dual metric, [18, 19] and [20].

The operator  $\Delta_g$  extends to a self-adjoint operator on  $L^2(M) \supset H^2(M) \rightarrow L^2(M)$  with compact resolvent. This implies that there exists an orthonormal basis  $\{\psi_k\} \in L^2(M)$  consisting of eigenfunctions such that

$$\Delta_g \psi_k = \lambda_k \psi_k \tag{2.3}$$

where the eigenvalues are listed with multiplicities

$$0 = \lambda_0 < \lambda_1 \le \lambda_2 \le \lambda_3 \le \dots \le \lambda_k \le \dots;$$
(2.4)

see for example, ([18,20] and [17]). The Laplacian  $\Delta_g$  thus, has one-dimensional null space consisting precisely of constant functions.

The *n*-dimensional sphere of radius  $r, r \in \mathbb{R}^+$  is the set of points in  $\mathbb{R}^{n+1}$  at a distance r from a given central point; i.e  $S^n(r) = \{x \in \mathbb{R}^{n+1} : ||x|| = r\}$ , [17,21,22] and [2]. We call  $S^n$  a unit *n*-sphere or simply an *n*-sphere when the radius r = 1 and write the unit *n*-sphere as the set

$$S^{n} = \{ x \in \mathsf{R}^{n+1} : || x || = 1 \}$$
(2.5)

where  $\|\cdot\|$  is the usual norm on  $L^2(S^n)$ , see e.g. [23].

The *n*-sphere is an *n*-dimensional compact manifold in (n+1)-space of constant positive sectional curvature, namely +1,  $n \ge 2$ . So, in particular, the 0-sphere, 1-sphere and the 2-sphere are respectively a pair of points on a line segment, a circle on a plane and the ordinary sphere in 3-dimension.

Let  $f \in S^n$  be any function on the *n*-sphere and  $\tilde{f}$  be its extension to an open neighbourhood of  $S^n$  that is constant along rays from the centre of  $S^n$ . We say that  $f \in C^2(S^n)$  if  $\tilde{f}$  is a  $C^2$  function of that neighbourhood. For such functions (not containing  $\{0\}$ ) on  $S^n$  the Laplacian  $\Delta_n$  equals

$$\Delta_n f = \Delta_g \tilde{f} \tag{2.6}$$

where  $\Delta_g$  on the right-hand side of (2.6) is the usual Laplacian in  $\mathbb{R}^{n+1}$ .

In  $\mathbb{R}^n$ ,  $n \ge 2$ , every point  $x \ne 0$  can be represented in polar coordinates as a couple  $(r, \theta)$  where r := |x| > 0 is the polar radius and  $\theta := \frac{x}{|x|} \in S^{n-1}$  is the polar angle. In the polar coordinates, the Riemannian measure on  $S^n$  is given by  $dV = \sin^{n-1} r dr d\theta$ .

For the unit n-sphere, the Laplacian (2.6) in polar coordinates reduces to

$$\Delta_{n} = \frac{1}{\sin^{n-1}\theta} \frac{\partial}{\partial \theta} \{\sin^{n-1}\theta \frac{\partial}{\partial \theta}\} + \frac{1}{\sin^{2}\theta} \Delta_{n-1}$$
(2.7)

where  $\Delta_{n-1}$  is the Laplacian on  $S^{n-1}$ .

Following [24, 14, 22] and [23], we give the following definitions.

**Definition 2.1** A function  $f : \mathbb{R}^n \to \mathbb{C}$  is called homogeneous of degree k if it satisfies  $f(tx) = t^k f(x)$  for all  $x \in \mathbb{R}^n$  and t > 0 fixed.

**Definition 2.2** Let  $\mathsf{P}_k(n)$  denote the space of homogeneous polynomials of degree k in (n+1) variables. The space  $\mathsf{H}_k(n) := \{ p_k \in \mathsf{P}_k(n) : \Delta_g p_k = 0, p_k \text{ homogeneous} \}$  is called the space of harmonic homogeneous polynomials.

Note, if  $p_k \in \mathsf{H}_k(n)$  then

$$p_k(x) = |x|^k \cdot p_k(\frac{x}{|x|})$$
; where  $\frac{x}{|x|} \in S^n$ ;  $x \neq 0$ .

Also, if  $p_k |_{S^n} = q_k |_{S^n}$ ; i.e  $p_k(x) = q_k(x) \quad \forall x \in S^n$ , then

$$p_k(x) = |x|^k p_k(\frac{x}{|x|}) = |x|^k q_k(\frac{x}{|x|}) = q_k(x) \quad \forall x \neq 0; \Longrightarrow p_k = q_k$$

since they are both polynomials. One may see [2, 25, 5] for more details.

**Theorem 2.3** [23]. The dimension  $d_k(n)$  of the space of harmonic polynomial  $H_k$  is given by the formula

$$d_k(n) = \binom{k+n}{n} - \binom{k+n-2}{n}.$$
(2.8)

We make the following lemma.

**Lemma 2.4** The multiplicities  $d_k(n)$  of the eigenspace of the spectrum  $\{\lambda_k\}$  can be expressed as

$$d_k(n) = \frac{(2k+n-1)(k+n-2)!}{k!(n-1)!}$$
(2.9)

where  $k \in N_0$  and  $n \ge 1$  is the dimension of the manifold  $S^n$ .

Proof. It is clear that

$$d_{k}(n) = \binom{k+n}{n} - \binom{k+n-2}{n} = \frac{(k+n)!}{k!n!} - \frac{(k+n-2)!}{(k-2)!n!}$$
$$= \frac{(k+n-2)!}{n!} [\frac{(k+n)(k+n-1)}{k!} - \frac{1}{(k-2)!}]$$

which simplifies as

$$\frac{(k+n-2)!}{n!(k-2)!} [\frac{(k+n)(k+n-1)}{k(k-1)} -1]$$
  
=  $\frac{(k+n-2)!}{k!} \frac{n}{n!} (2k+n-1)$   
=  $\frac{(2k+n-1)(k+n-2)!}{k!(n-1)!}$ .

#### 2.2 Spectral zeta function

Spectral zeta function is best explained through the well-known Riemann zeta function. Recall that the Riemann zeta function  $\zeta_R$  is the function defined as  $\zeta_R : \{s \in \mathbb{C} : \Re(s) > 1\} \rightarrow \mathbb{C}$  with

$$\zeta_{R}(s) = \sum_{k=1}^{\infty} \frac{1}{k^{s}}; \ \Re(s) > 1$$
(2.10)

c.f: [4] and [10]. Notice that since

$$\sum_{k=1}^{\infty} \left| \frac{1}{k^s} \right| = \sum_{k=1}^{\infty} \frac{1}{k^{\Re(s)}},\tag{2.11}$$

the series on the right-hand-side of (2.11) converges absolutely if and only if  $\Re(s) > 1$ . The Riemann zeta function defined by (2.10) above is holomorphic in the region indicated. It, however, admits a meromorphic

continuation to the whole s-complex plane with only simple pole at s = 1 and has residue 1; see e.g. [19,21, 16, 4, 10].

A generalization of the Riemann zeta function (2.10) is the Hurwitz zeta function  $\zeta_H(s, a)$ .

**Definition 2.5** [4,10] Let  $s \in \mathbb{C}$  and  $0 < a \leq 1$ . Then for  $\Re(s) > 1$ , the Hurwitz zeta function is defined by

$$\zeta_{H}(s,a) = \sum_{k=0}^{\infty} \frac{1}{(k+a)^{s}}; \ \Re(s) > 1.$$
(2.12)

Clearly,  $\zeta_H(s,1) = \zeta_R(s)$ . Expression for  $a = b+1; b \in \mathbb{R}$  follows by observing that

$$\zeta_{H}(s,1+b) = \sum_{k=0}^{\infty} \frac{1}{(k+1+b)^{s}} = \zeta_{H}(s,b) - \frac{1}{b^{s}}.$$
(2.13)

**Theorem 2.6** For  $0 < a \le 1$ , we have

$$\zeta_{H}(s,a) = \frac{1}{a^{s}} + \sum_{m=0}^{\infty} (-1)^{m} \frac{\Gamma(s+m)}{m! \Gamma(s)} a^{m} \zeta_{R}(s+m).$$
(2.14)

*Proof.* Note that for |z| < 1 the following binomial expansion holds

$$(1-z)^{-s} = \sum_{m=0}^{\infty} \frac{\Gamma(s+m)}{m! \Gamma(s)} z^m.$$

So for  $\Re(s) > 1$ , we have

$$\zeta_{H}(s,a) = \frac{1}{a^{s}} + \sum_{k=1}^{\infty} \frac{1}{k^{s}} \frac{1}{(1 + \frac{a}{k})^{s}}$$
$$= \frac{1}{a^{s}} + \sum_{k=1}^{\infty} \frac{1}{k^{s}} \sum_{m=0}^{\infty} (-1)^{m} \frac{\Gamma(s+m)}{m!\Gamma(s)} (\frac{a}{k})^{m}$$
$$= \frac{1}{a^{s}} + \sum_{m=0}^{\infty} (-1)^{m} \frac{\Gamma(s+m)}{m!\Gamma(s)} (a)^{m} \sum_{k=1}^{\infty} \frac{1}{k^{s+m}}$$

which gives the expansion

Another generalisation of the Riemann zeta function is the spectral zeta function, which is the function of interest in this paper. The Laplacian defined in (2.2) acting on smooth functions on the closed and connected Riemannian *n*-dimensional manifolds is a non-negative operator and has the discrete spectrum  $\{\lambda_k\}_{k=1}^{\infty}$  listed with multiplicities. We define

$$\zeta_{g}(s) = \sum_{k=1}^{\infty} \frac{1}{\lambda_{k}^{s}}; \quad \Re(s) > \frac{n}{2};$$

$$(2.15)$$

see e.g. [21,11].

The integral kernel,  $\zeta_g(s, x, y)$ , associated with the spectral zeta function is defined via the operator  $\Delta_g^{-s}$  (see e.g. [9]) given by the following properties, [23] and [6].

- 1. It is linear on  $L^2(M)$  with 1-dimensional null space consisting of constant functions. This ensures that the smallest eigenvalue of  $\Delta_g^{-s}$  is 0 of multiplicity 1 with corresponding eigenfunction  $\frac{1}{\sqrt{V}}$  where V is the volume of M.
- 2. The image of  $\Delta_g^{-s}$  is contained in the orthogonal complement of constant functions in  $L^2(M)$  i.e.

$$\int_{M} \Delta_{g}^{-s} \psi dV_{g} = 0 \ \forall \psi \in L^{2}(M) \text{ constant.}$$

3.  $\Delta_g^{-s} \psi_k(x) = \lambda_k^{-s} \psi_k(x)$  for all  $\psi_k$ ; k > 0 an orthonormal basis of eigenfunction of  $\Delta_g$ .

Then for  $\Re(s) > \frac{n}{2}$ , we see by property (3.) that  $\Delta_g^{-s}$  is trace class, with trace given by the spectral zeta function, namely

$$\zeta_{g}(s) = \sum_{k=1}^{\infty} \frac{1}{\lambda_{k}^{s}} = \operatorname{Tr}(\Delta_{g}^{-s}) = \int_{M} \zeta_{g}(s, x, x) dV ; \quad \Re(s) > \frac{n}{2}.$$
(2.16)

**Theorem 2.7** [17,25]. Let  $\{\psi_k\}_{k=1}^{\infty}$  be an orthonormal eigenbasis for  $\Delta_g$  corresponding to the eigenvalues  $\{\lambda_k\}_{k=1}^{\infty}$  listed with multiplicities. Then the zeta kernel,  $\zeta_g(s, x, y)$ , also called the **point-wise zeta** function, is equal to

$$\zeta_{g}(s,x,y) = \sum_{k=1}^{\infty} \frac{\psi_{k}(x)\overline{\psi}_{k}(y)}{\lambda_{k}^{s}}; \ \Re(s) > \frac{n}{2}.$$
(2.17)

From here on, we suppress the subscript g in  $\zeta_g(s)$  and  $\Delta_g$ . We simply write  $\zeta(s)$  and  $\Delta$  for  $\zeta_g(s)$  and  $\Delta_g$  respectively, unless for purpose of emphasis.

## **3** Casimir Energy of $\Delta_g$ on $S^n$

Using the properties of the Riemann and Hurwitz zeta functions reviewed above, we can now compute the Casimir energy of the Laplacian  $\Delta_g$  for the round metric on  $S^n$ . As mentioned earlier, sometimes zeta

regularisation is not sufficient to define the Casimir energy, because  $\zeta_g$  has a pole at -1/2. In this case, we define the Casimir energy as the finite part of  $\zeta_g$  at s = -1/2, where FP is the finite part function given by

$$FP[f](s) := \begin{cases} f(s) \text{ if } s \text{ is not a pole} \\ \lim_{\varepsilon \to 0} (f(s+\varepsilon) - \frac{\text{Residue}}{\varepsilon}), \text{ if } s \text{ is a pole.} \end{cases}$$
(3.1)

Now note that the zeta function on the n-sphere is specifically given by

$$\zeta_{s^n}(s) = \sum_{k=1}^{\infty} \frac{d_k(n)}{(k(k+n-1))^s}; \quad \Re(s) > \frac{n}{2}$$
(3.2)

with  $d_k(n)$  defined by equation (2.8), [2,13,12,17,25,23,5].

The following theorems will be used to switch limits and integrals; and sum and integrals.

**Theorem 3.1** [18,19]. (Dominated Convergence Theorem). Let  $\Omega \subseteq \mathbb{R}^n$  be open and let  $\{\psi_k\}$  be a sequence of integrable functions on  $\Omega$ . Suppose that  $\lim_{k\to\infty} \psi_k(x) = \psi(x)$   $\mu$  -almost everywhere. Further suppose that there exists  $\omega \ge 0$  with  $\int_{\Omega} \omega(x) d\mu(x) < \infty$  such that  $\psi_k(x) \le \omega(x) \quad \forall k$ . Then  $\psi(x) \le \omega(x) \quad \mu$  -almost everywhere and

$$\lim_{k\to\infty}\int_{\Omega}\psi_k(x)\mathrm{d}\mu(x)=\int_{\Omega}\psi(x)\mathrm{d}\mu(x);$$

where  $d\mu(x)$  is the measure form on  $\Omega$ .

**Theorem 3.2** [18,19]. (Fubini - Tonelli theorem). Let  $\{\Psi_k\}$  be a sequence of measurable functions. Sum and integral such as  $\sum_k \int \Psi_k(x) dx$  can be interchanged in either of the following cases:

$$\Psi_k \ge 0, \forall k \in \mathbb{N} \text{ or } \sum_k \int |\Psi_k(x)| \, dx < \infty$$

Thus  $\Re(s) > \frac{n}{2}$  one writes  $\zeta_{s^n}(s)$  from (3.2) as

$$\zeta_{s^n}(s) = \frac{1}{(n-1)!} \sum_{k=1}^{\infty} \frac{(k+n-2)!(2k+n-1)}{k!(k(k+n-1))^s}$$

and make the substitution

$$w = k + \frac{n-1}{2} \tag{3.3}$$

to have

$$\zeta_{s^n}(s) = \sum_{w=1}^{\infty} \frac{2w\Gamma(w + \frac{n-1}{2})}{\Gamma(n)\Gamma(w + \frac{3-n}{2})(w^2 - \frac{(n-1)^2}{4})^s}.$$

By Mellin transform

$$\zeta_{s^{n}}(s) = \frac{1}{\Gamma(s)} \int_{0}^{\infty} \sum_{w=1}^{\infty} \frac{2w\Gamma(w + \frac{n-1}{2})}{\Gamma(n)\Gamma(w + \frac{3-n}{2})} e^{-w^{2}t} e^{\frac{(n-1)^{2}}{4}t} t^{s-1} dt.$$

Re-arranging of integral and sum using the Fubini - Tonelli theorem (3.2) since the exponential of a negative number is bounded gives

$$\zeta_{s^{n}}(s) = \frac{1}{\Gamma(s)} \sum_{m=0}^{\infty} \sum_{w=1}^{\infty} \frac{(n-1)^{2m}}{m! 4^{m}} \frac{2w\Gamma(w+\frac{n-1}{2})}{\Gamma(n)\Gamma(w+\frac{3-n}{2})} \int_{0}^{\infty} e^{-w^{2}t} t^{s+m-1} dt.$$

Now let  $w^2 t \mapsto \tau$  to have

$$\begin{aligned} \zeta_{s^{n}}(s) &= \frac{1}{\Gamma(s)} \sum_{m=0}^{\infty} \sum_{w=1}^{\infty} \frac{(n-1)^{2m}}{m! 4^{m}} \frac{2w\Gamma(w+\frac{n-1}{2})}{\Gamma(n)\Gamma(w+\frac{3-n}{2})} w^{-2s-2m} \int_{0}^{\infty} e^{-\tau} \tau^{s+m-1} d\tau \\ &= 2 \sum_{m=0}^{\infty} \frac{(n-1)^{2m}}{m! 4^{m}} \frac{\Gamma(s+m)}{\Gamma(s)} \sum_{w=1}^{\infty} \frac{w^{-2s-2m+1}\Gamma(w+\frac{n-1}{2})}{\Gamma(n)\Gamma(w+\frac{3-n}{2})}. \end{aligned}$$

Thus,

$$\zeta_{s^{n}}(s) = 2\sum_{m=0}^{\infty} \frac{(n-1)^{2m}}{m! 4^{m}} \frac{\Gamma(s+m)}{\Gamma(s)} \sum_{k=1}^{\infty} \frac{\Gamma(k+n-1)}{\Gamma(n)\Gamma(k+1)} (k+\frac{n-1}{2})^{-2s-2m+1}.$$
(3.4)

.

Consequently, for n = 1, ie the unit circle, we have

$$\zeta_{s^1}(s) = \sum_{k=1}^{\infty} \frac{2}{k^{2s}} = 2\zeta_R(2s)$$
(3.5)

which gives

$$\zeta_{s^1}(-\frac{1}{2}) = 2\zeta_R(-1) \approx -0.1666667.$$
(3.6)

For  $S^2$ , we have

$$\zeta_{s^2}(s) = \sum_{k=1}^{\infty} \frac{2k+1}{(k(k+1))^s}$$
(3.7)

and from 3.3 making the substitution  $w = k + \frac{1}{2}$  yields

$$\zeta_{s^2} = \sum_{w=1}^{\infty} \frac{2w}{(w^2 - \frac{1}{4})^s}$$

which reduces from (3.4) to

$$\zeta_{s^2} = 2\sum_{m=0}^{\infty} \frac{\Gamma(s+m)}{m! 4^m \Gamma(s)} \zeta_H (2s+2m-1,\frac{3}{2}).$$
(3.8)

Therefore, the Casimir energy  $\zeta_{s^2}(-\frac{1}{2})$  of  $\Delta_g$  on  $S^2$  becomes

$$\operatorname{FP}[\zeta_{s^2}(-\frac{1}{2})] \approx -0.265096. \tag{3.9}$$

Similarly, on the 3-sphere, we have

$$\zeta_{s^3}(s) = \sum_{k=1}^{\infty} \frac{(k+1)^2}{(k(k+2))^s}$$
(3.10)

and let w = k + 1 so that

$$\zeta_{s^3}(s) = \sum_{w=1}^{\infty} \frac{w^2}{(w^2 - 1)^s}$$

and thus

$$\zeta_{s^3}(s) = \sum_{m=0}^{\infty} \frac{\Gamma(s+m)}{m! \Gamma(s)} \zeta_H(2s+2m-2,2)$$
(3.11)

which has simple pole at m = 2 with residue  $-\frac{1}{6}$  for  $\zeta_{s^3}(-\frac{1}{2})$ . Therefore,

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$$FP[\zeta_{s^3}(-\frac{1}{2})] \approx -0.411503. \tag{3.12}$$

On  $S^4$ , we have

$$\zeta_{s^4}(s) = \frac{1}{6} \sum_{k=1}^{\infty} \frac{(k+1)(k+2)(2k+3)}{(k(k+3))^s}$$
(3.13)

and let  $w = k + \frac{3}{2}$  to get

$$\zeta_{s^4}(s) = \sum_{w=1}^{\infty} \frac{w(w^2 - \frac{1}{4})}{(w^2 - \frac{9}{4})^s}.$$

From (3.4) this reduces to

$$\zeta_{s^4}(s) = \frac{1}{3} \sum_{m=0}^{\infty} \frac{\Gamma(s+m)}{m! \Gamma(s)} (\frac{9}{4})^m (\zeta_H(2s+2m-3,\frac{5}{2}) - \frac{1}{4} \zeta_H(2s+2m-1,\frac{5}{2})).$$
(3.14)

Unsurprisingly,  $s = -\frac{1}{2}$  is not a pole for  $\zeta_{s^4}(s)$ . This gives the Casimir energy of the Laplacian on  $S^4$  as

$$\zeta_{s^4}(-\frac{1}{2}) = \frac{1}{3} \sum_{m=0}^{\infty} \frac{\Gamma(m-1/2)}{m! \Gamma(-1/2)} (\frac{9}{4})^m (\zeta_H (2m-4, \frac{5}{2}) - \frac{1}{4} \zeta_H (2m-2, \frac{5}{2}))$$
  

$$\approx -0.431743.$$
(3.15)

Furthermore, for  $S^5$ , we have

$$\zeta_{s^5}(s) = \frac{1}{12} \sum_{k=1}^{\infty} \frac{(k+1)(k+2)^2(k+3)}{(k(k+4))^s}.$$
(3.16)

Let w = k + 2 to get

$$\zeta_{s^5}(s) = \frac{1}{12} \sum_{w=1}^{\infty} \frac{w^2(w^2 - 1)}{(w^2 - 4)^s}.$$

Again from (3.4) this reduces to

$$\zeta_{s^{5}}(s) = \frac{1}{12} \sum_{m=0}^{\infty} \frac{4^{m} \Gamma(s+m)}{m! \Gamma(s)} (\zeta_{H}(2s+2m-4,3) - \zeta_{H}(2s+2m-2,3)).$$
(3.17)

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The formula (3.17) has simple poles for  $s = -\frac{1}{2}$  at m = 2 and m = 3 with residues  $\frac{1}{12}$  and  $-\frac{1}{6}$  respectively. Hence, the Casimir energy of the Laplacian on  $S^5$  is

$$FP[\zeta_{s^5}(-\frac{1}{2})] = -0.510570. \tag{3.18}$$

The Casimir energy  $\zeta_{\Delta_g}(-\frac{1}{2})$  of the Laplacian  $\Delta_g$  on higher dimensional unit spheres may be computed similarly using the formula (3.4).

## 4 Conclusion

We have shown that for a given dimension n of the unit sphere, our formula (3.4) reduces to a simpler formula in terms of the Riemann zeta function or its generalisation, the Hurwitz zeta function. The Casimir energy is then simply read-off as

$$FP[\zeta_{s^n}(s)]|_{s=\frac{1}{2}} = FP[2\sum_{m=0}^{\infty} \frac{(n-1)^{2m}}{m!4^m} \frac{\Gamma(s+m)}{\Gamma(s)} \sum_{k=1}^{\infty} \frac{\Gamma(k+n-1)}{\Gamma(n)\Gamma(k+1)} (k+\frac{n-1}{2})^{-2s-2m+1}]|_{s=\frac{1}{2}}.$$

This approach may be employed to compute many other special values of the spectral zeta function of the Laplacian on some other Riemannian manifolds.

## **Competing Interests**

Author has declared that no competing interests exist.

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